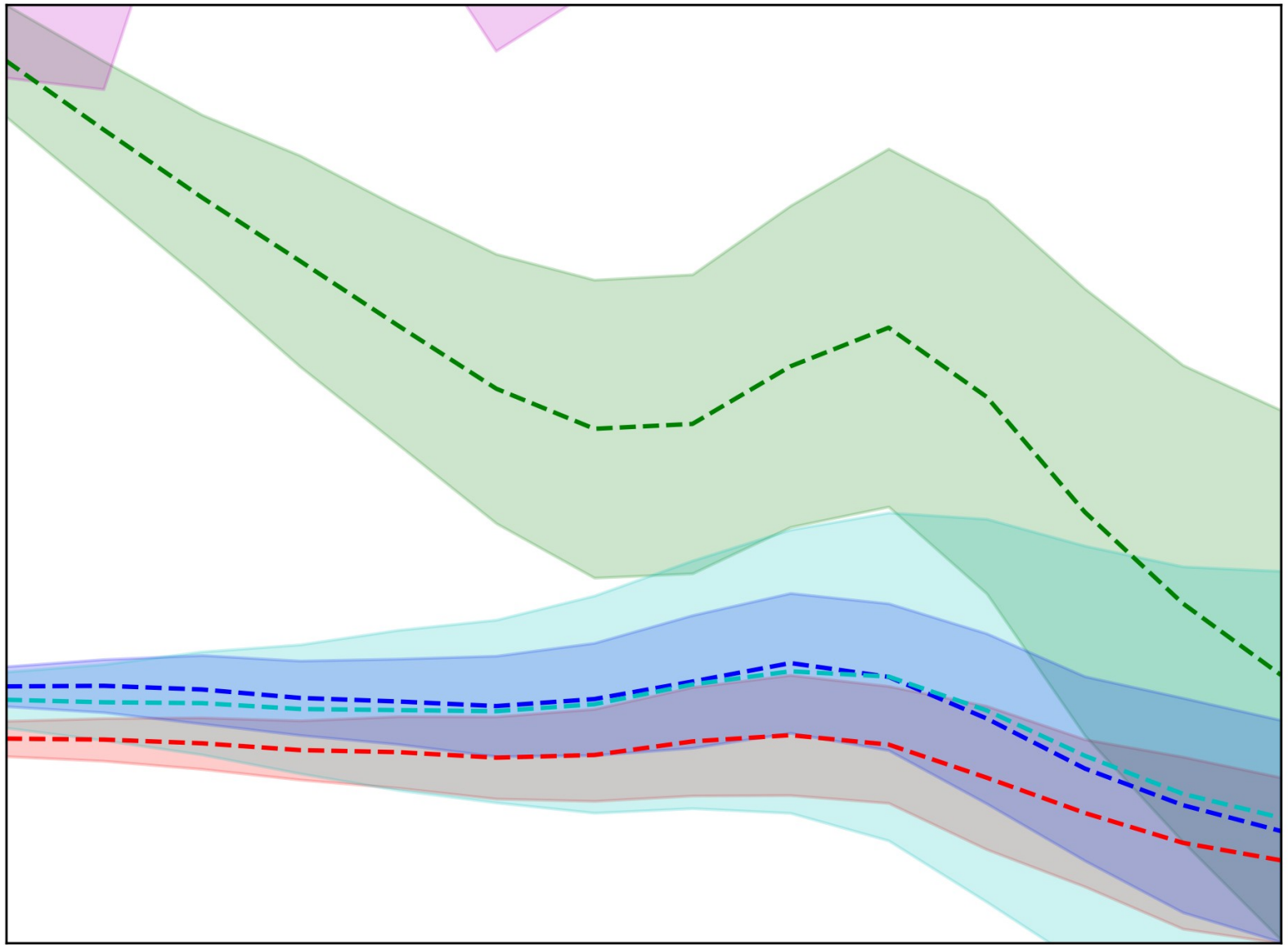


Using cross-correlations to recover the BAO peak in sparse spectroscopic surveys

by **Christoph Saulder** (KASI)

8th Korea-Japan workshop on Dark Energy



Collaborators

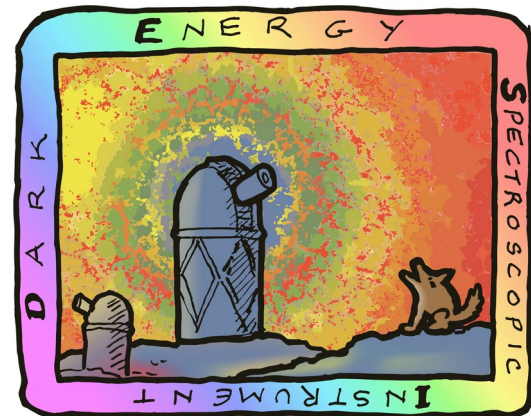
- Christoph Saulder (KASI)
- Yong-Seon Song (KASI)
- Minji Oh (formerly SJTU)
- Feng Shi (Xidian University)
- Yi Zheng (SYSU)



한국천문연구원
Korea Astronomy & Space Science Institute



西安电子科技大学
XIDIAN UNIVERSITY

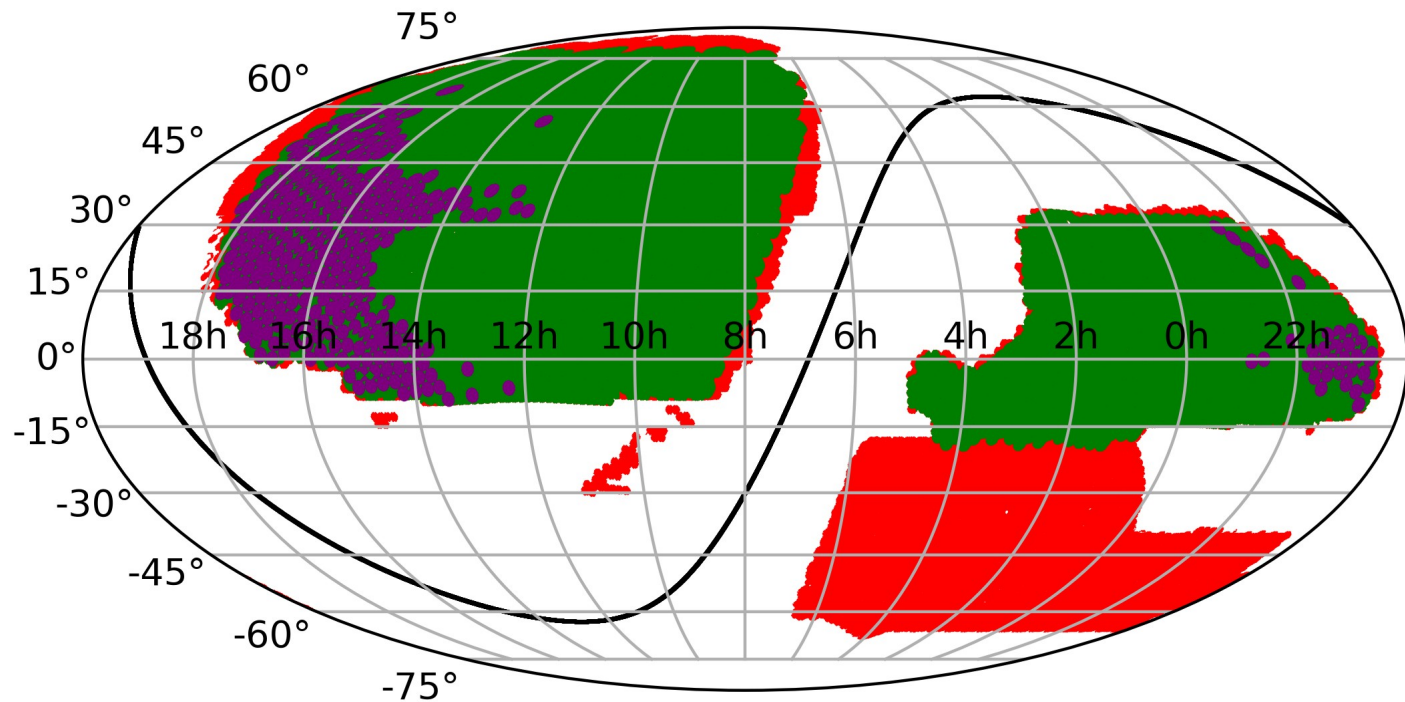


The Motivation

- Photometric survey get completed before spectroscopic surveys
- Spectroscopic surveys will usually be less complete than photometric surveys
- However photometric redshifts have huge error bars
- Depending on survey strategy for spectroscopic survey: first wide than deep or first deep than wide

DESI DA0.2 footprint

- DESI photometric footprint
- DESI spectroscopic footprint
- DESI spectroscopy by July 2021
- galactic equator



Our toy model

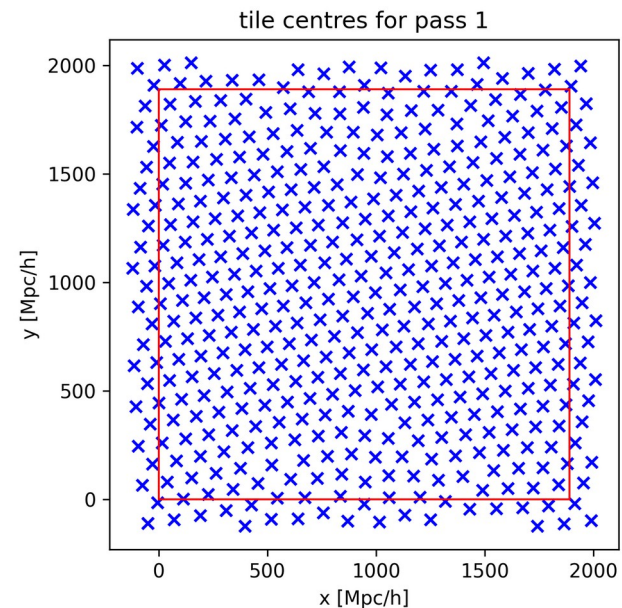
- DESI-like survey
- Photo-z uncertainty $\sigma_0 = 0.02$
- Current focus: LRGs around $z=0.7$
- Only one target class
- Fibre assignment based on DESI

Simulations

- 100 mocks with Planck cosmology
- Cubes with 1890 Mpc/h side length
- Halos identified using Rockstar
- Galaxies are painted using HOD models based on observations (currently only tested with the DESI LRG sample)

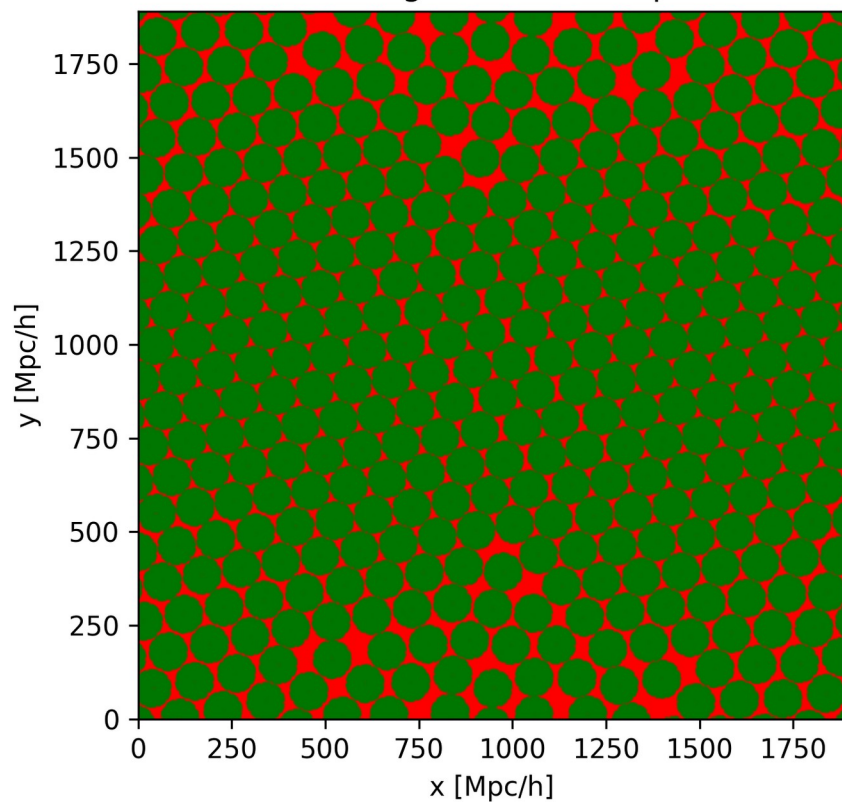
Fibre assignment

- Tiling strategy of DESI
- Comparing 1-pass with many passes
- Impact of fibre placements and completeness of the spectroscopic data

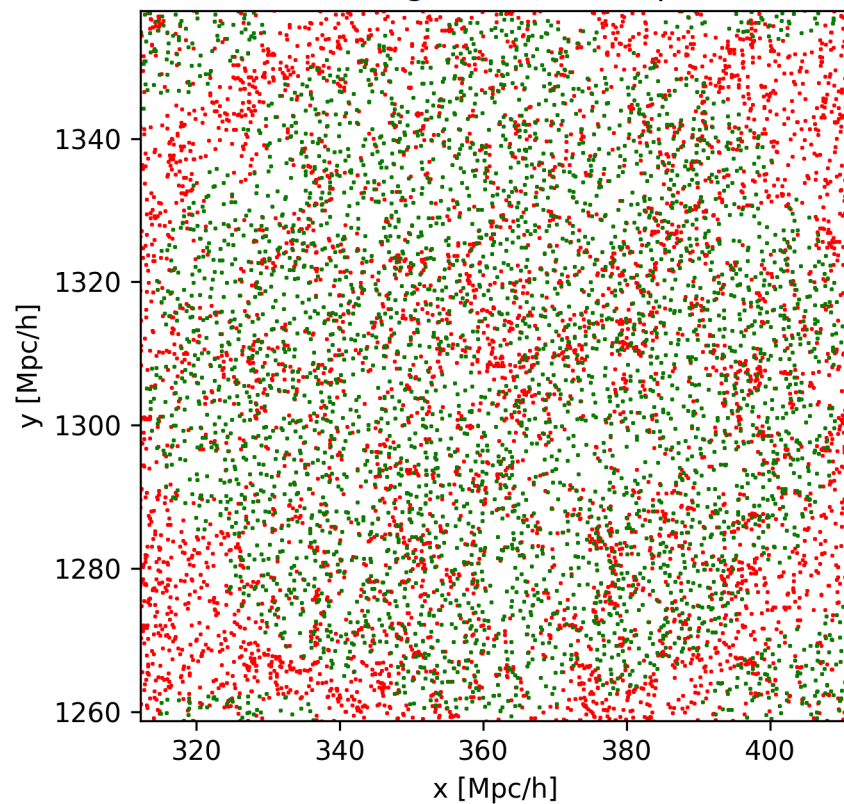


First pass

fibre assignment after 1 pass



fibre assignment after 1 pass

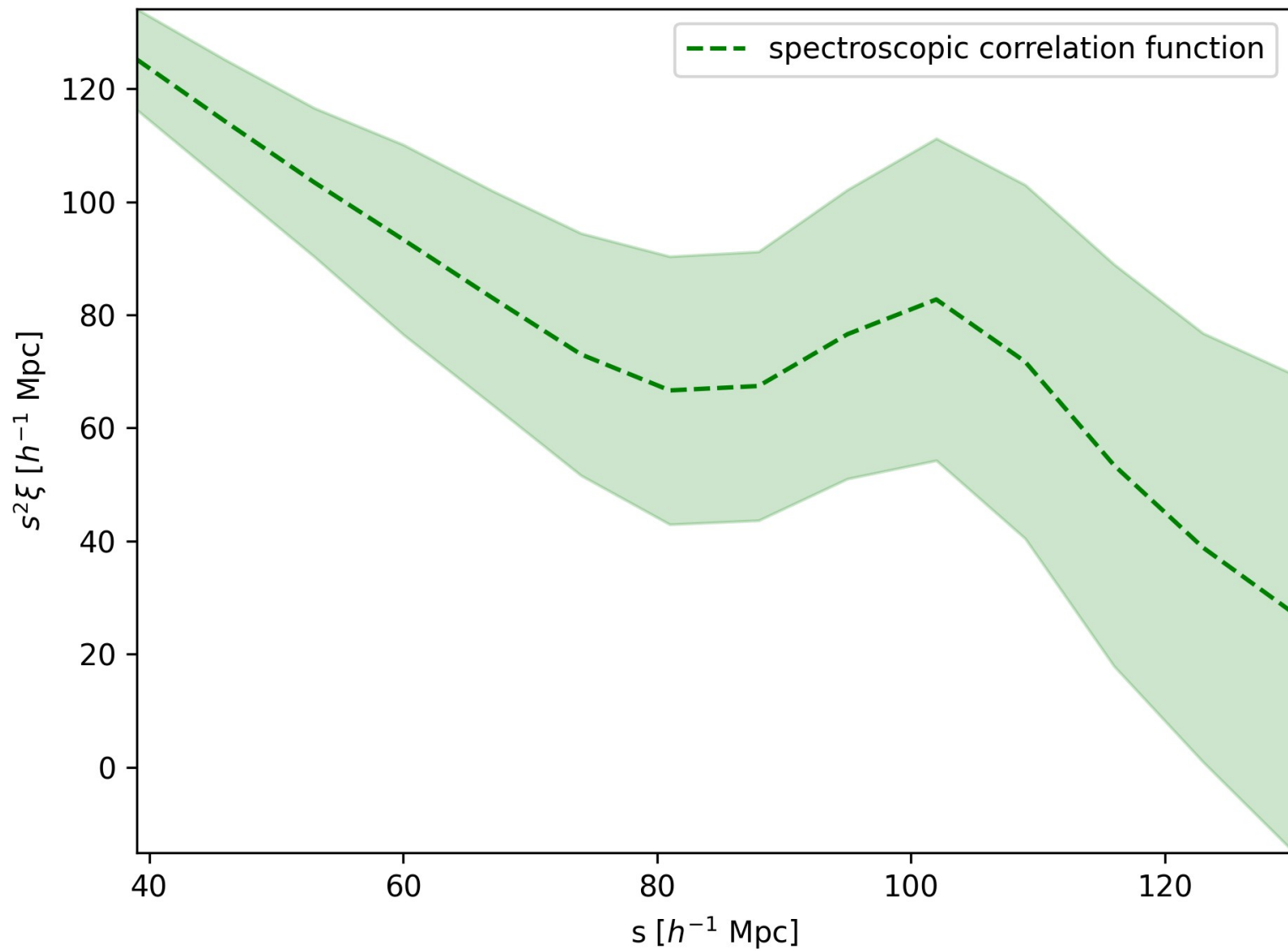


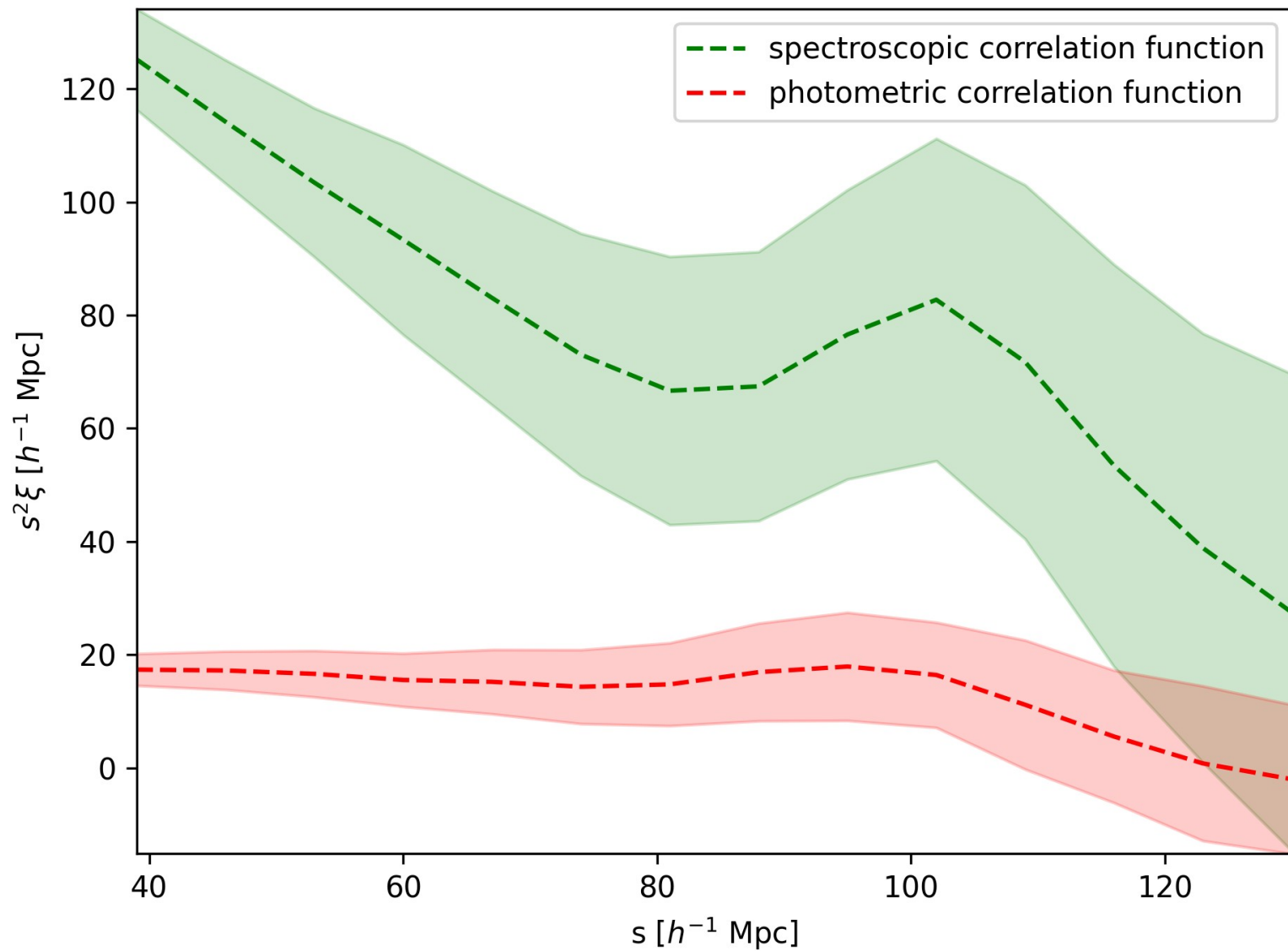
First pass

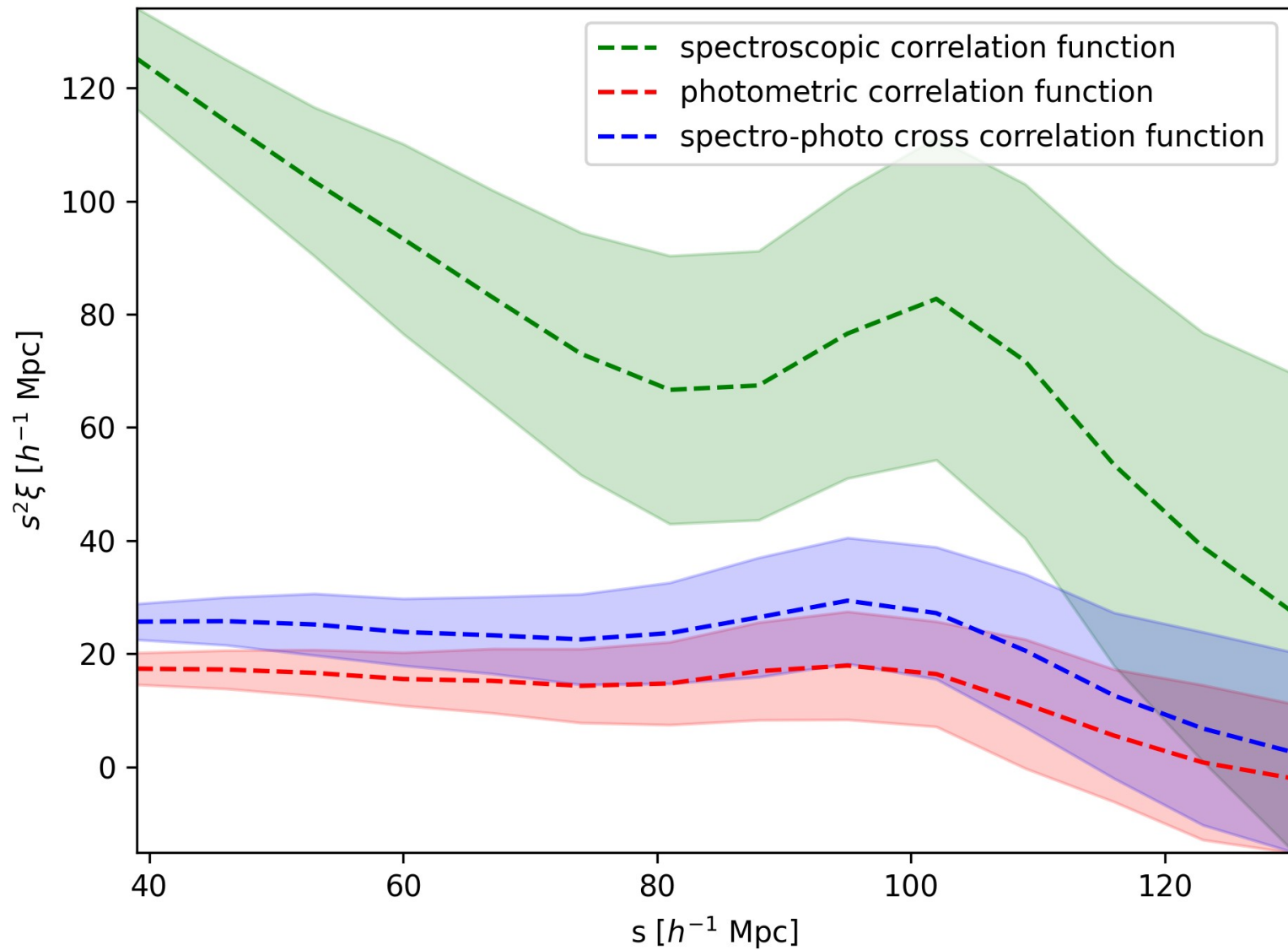
- Comparable to the current status of DESI and the first DESI main survey data release
- In simulation: about 40% of targets covered
- In observations: about 20% completeness after 1 pass for LRG, because broken fibres, sky and star fibres, competing other targets, ...

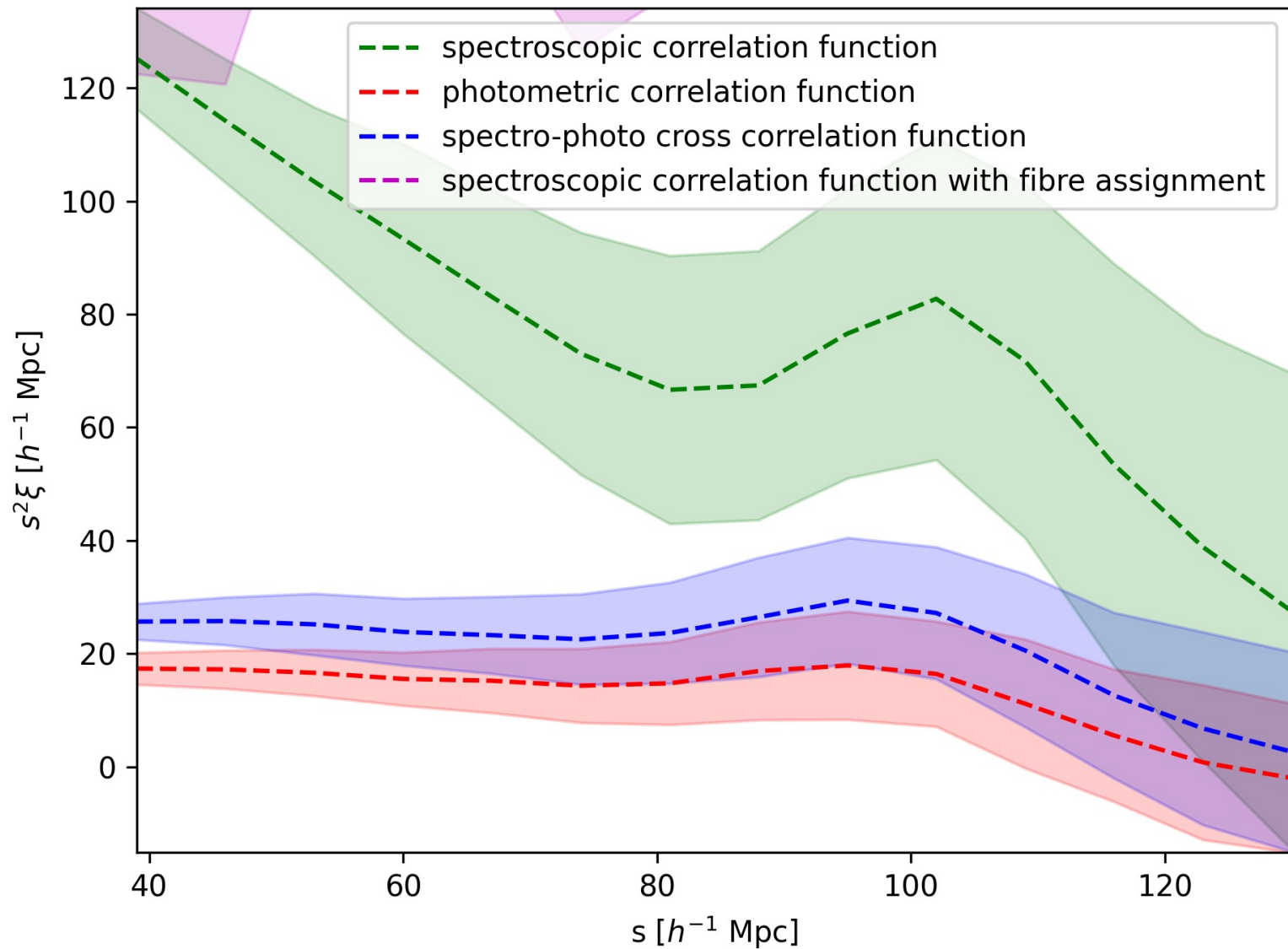
Large scale correlation function

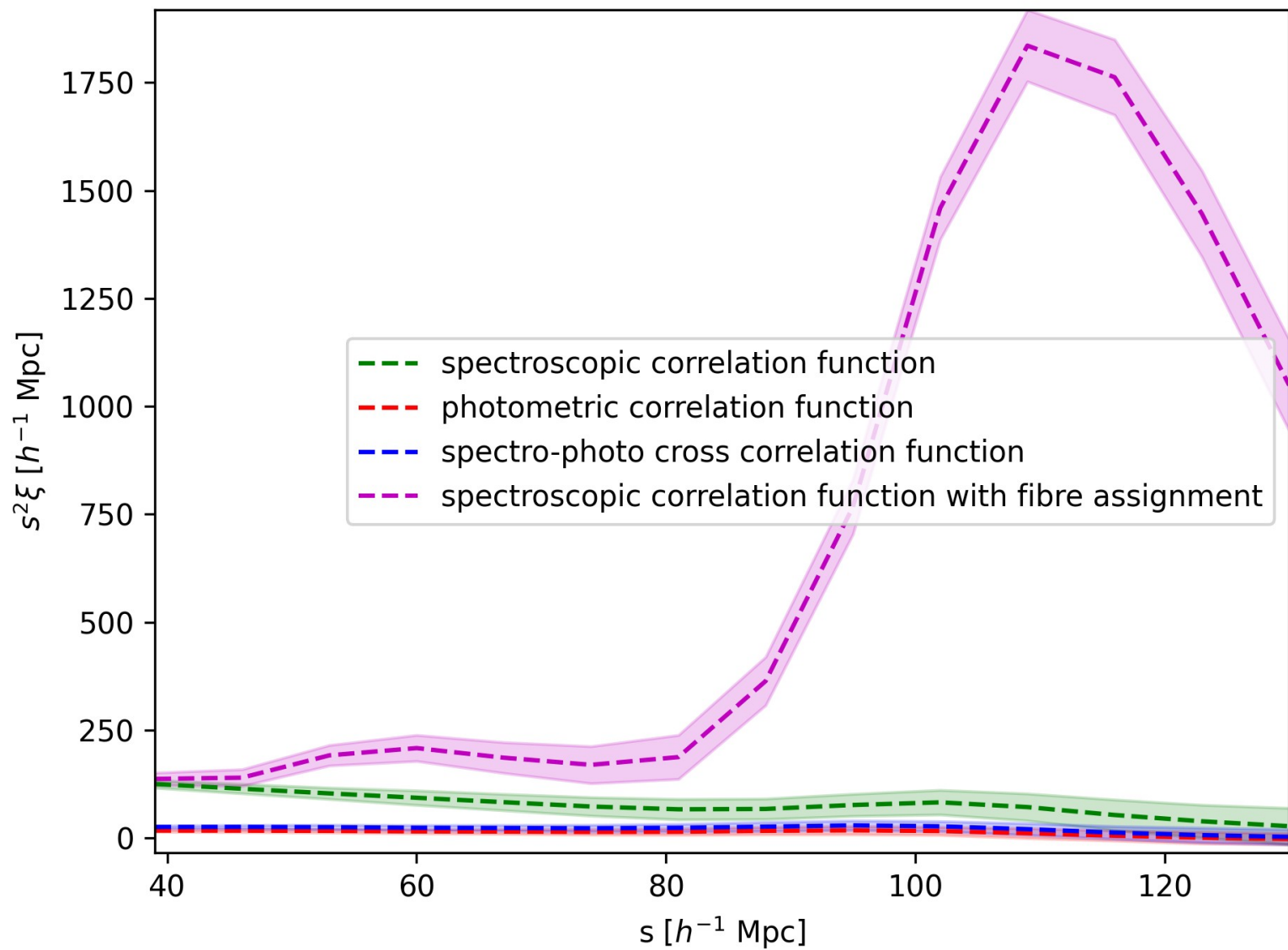
- Correlation function $\xi(s, \mu)$ around BAO peak
- Comparison between
 - Clean spectroscopic data
 - Photometric data
 - Cross-correlation spectro-photo data
 - Spectroscopic data after fibre assignment

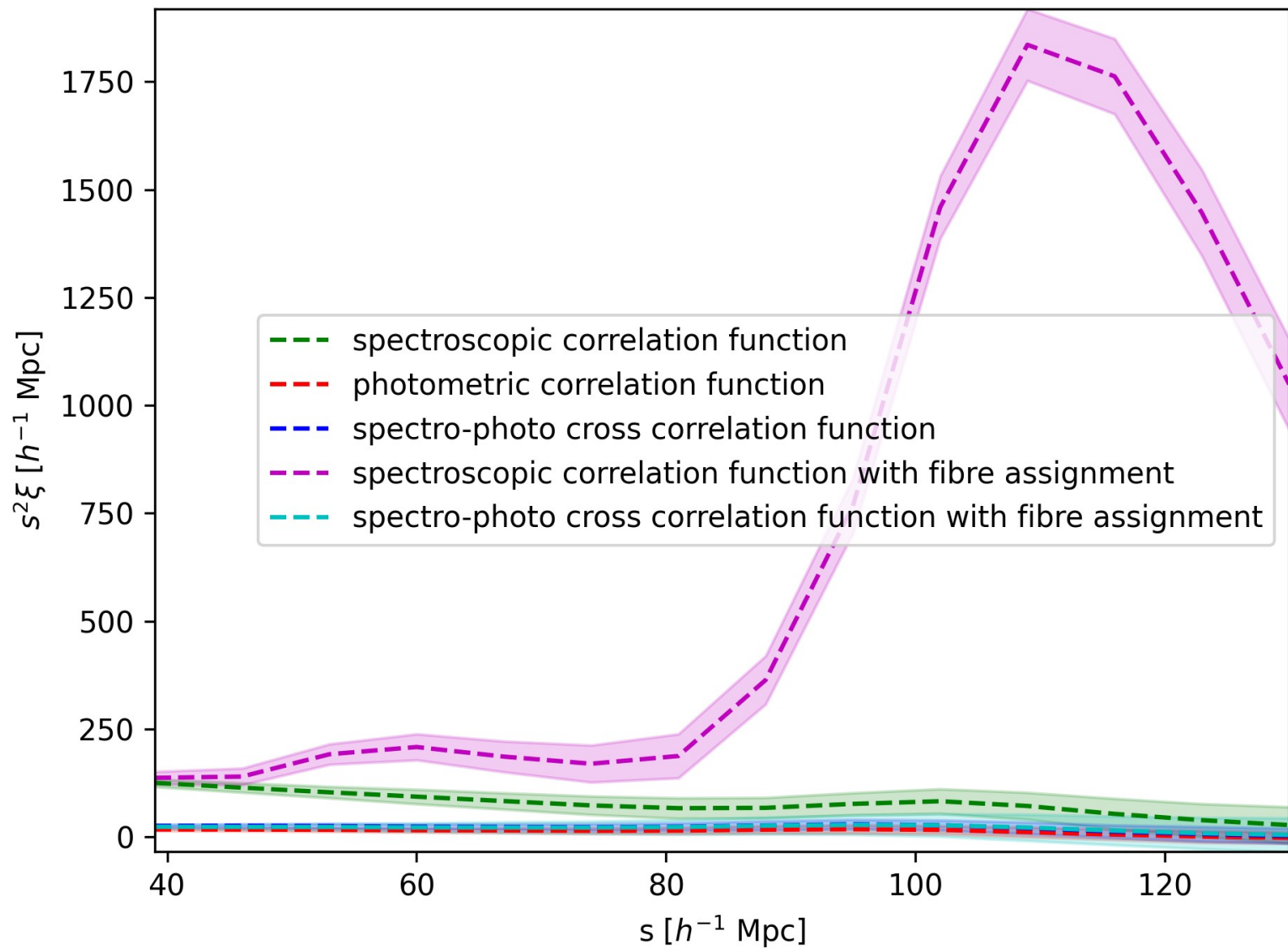


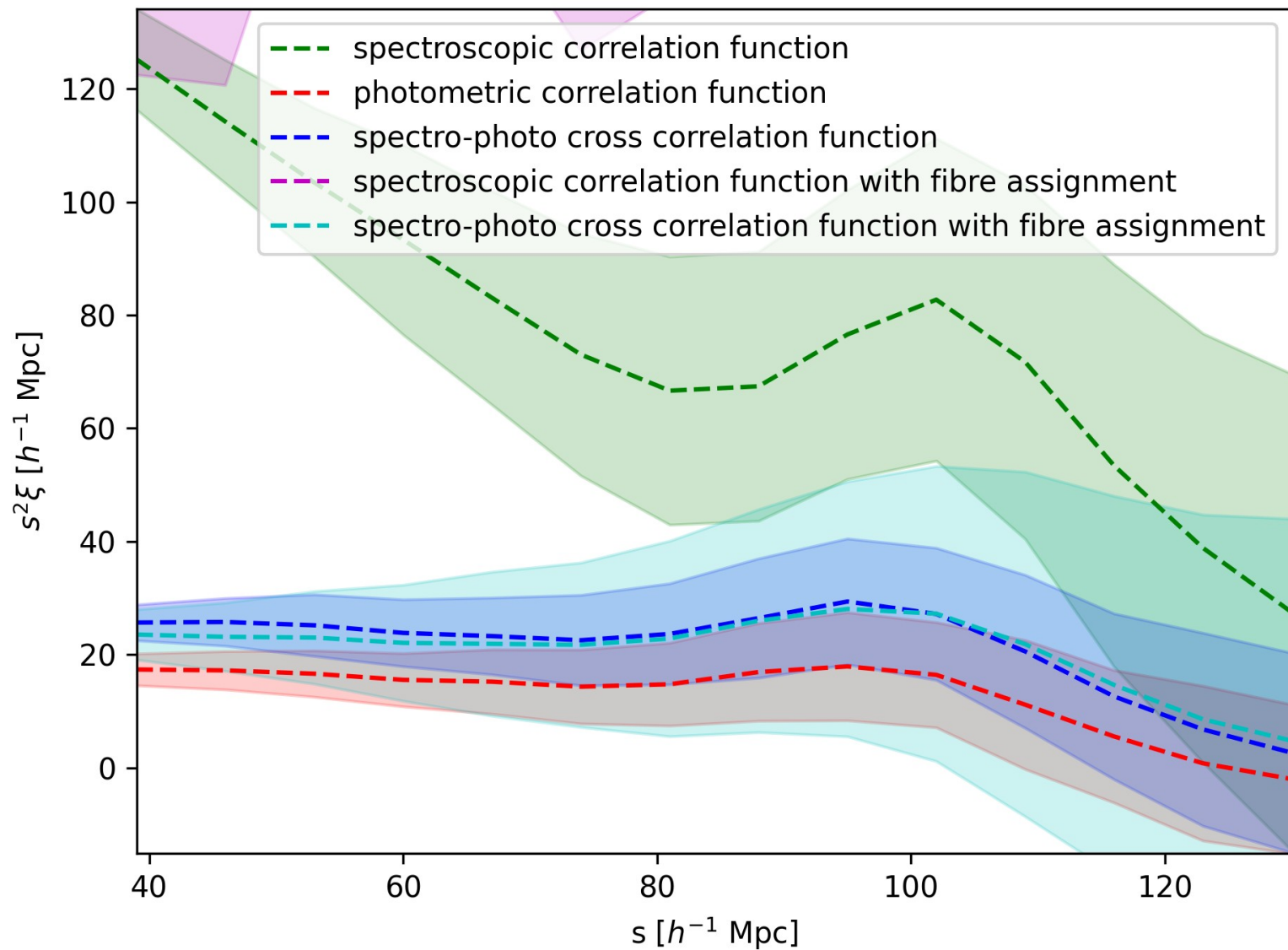












Outlook

- More sophisticated toy models
- Considering possible other target classes
- Testing if a possible competitive probe to other techniques (such as PIP weights) for early DESI data
- Considering other future surveys (J-PAS, Euclid, LSST)

Summary

- Comparing auto- and cross-correlation function for photometric and spectroscopic LRG selection
- Cross-correlation can recover features, even if the spectroscopic data set is incomplete and biased due to fibre assignment
- Application for early DESI data (DA0.2)

QUESTIONS?

Thank you!
감사합니다!
ありがとうございました!

